

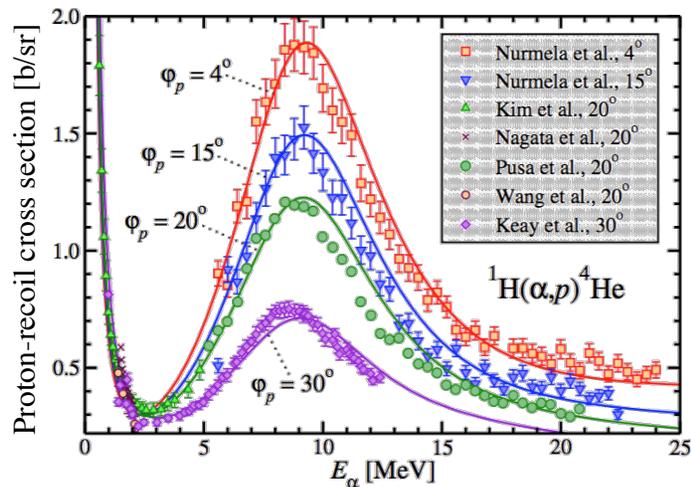
Fundamental theory of scattering and reactions

Developing a fundamental theory of nuclear scattering and reactions is crucial for nuclear science, nuclear astrophysics, and applications

- Exotic nuclei offer an exciting opportunity to test our understanding of nuclear properties in terms of forces emerging from QCD, but:
 - Need theory able to accurately model the reactions used to populate ground states and resonances of exotic nuclei
 - Only with a unified treatment of bound and continuum states can we correctly interpret, e.g., the role of the 3N force in exotic nuclei
- Predictive theory needed at low energies beyond experimental reach, to improve evaluation of important rates for Big Bang Nucleosynthesis, Standard Solar Model, Stellar Nucleosynthesis, fusion energy research

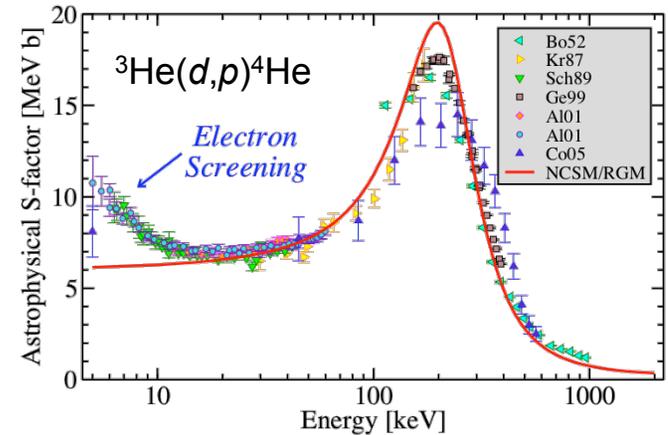
Present status

- *Ab initio* approach with realistic NN forces developed and applied to two-body light-nucleus reactions relevant for astrophysics and fusion research
- *First ab initio* description of ${}^4\text{He}+n+n$ continuum paves the way to fusion reactions with 3-body final states
- Complete *ab initio* descriptions of scattering in the $A=5$ sector give us initial confidence on the methods and underlying NN+NNN Hamiltonians
- *Ab initio* studies of N - and d -nucleus scattering give first glimpse of interplay between NNN force and continuum effects in energy spectra

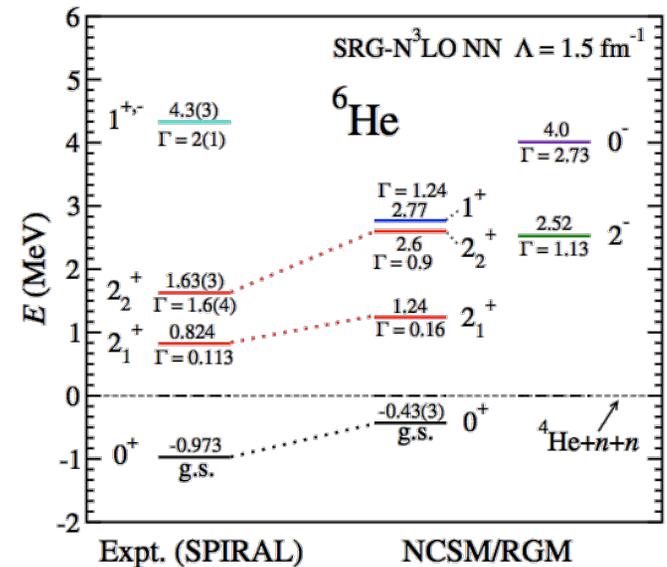


Left: complete *ab initio* description of proton- ${}^4\text{He}$ scattering

Right: ${}^6\text{He}$ spectrum of states from first *ab initio* description of ${}^4\text{He}+n+n$ continuum



$d+{}^3\text{He}$ fusion from *ab initio* theory with realistic NN interaction vs Expt.

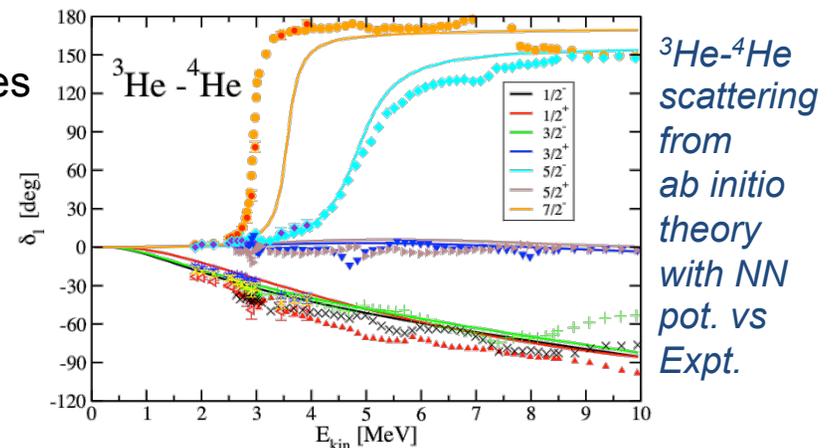


Future challenges

- Complete validation for our theory/NN+NNN Hamiltonians and deliver accurate predictions for lesser known $A=5$ cross sections
 - Stringent test on experimentally well-known ${}^3\text{H}(d,n){}^4\text{He}$ cross section
 - $d+{}^3\text{H}\rightarrow{}^4\text{He}+n+\gamma$ bremsstrahlung (fusion research); ${}^3\text{H}(d,p){}^4\text{He}$ (hindered by electron screening)

- Provide *ab initio* description of binary reactions involving p-shell nuclei and/or composite projectiles

- (d,p) transfer reactions relevant for astrophysics and FRIB (e.g.: ${}^7\text{Li}(d,p){}^8\text{Li}$)
- ${}^3\text{He}(\alpha,\gamma){}^7\text{Be}$, ${}^7\text{Be}(p,\gamma){}^8\text{Be}$ (standard solar model); ${}^6\text{Li}(n,{}^3\text{H}){}^4\text{He}$ (fusion research)
- ${}^8\text{Be}(\alpha,\gamma){}^{12}\text{C}$, ${}^{12}\text{C}(\alpha,\gamma){}^{16}\text{O}$ (stellar nucleosynthesis), ...



- Deliver high-fidelity simulations complete of both NNN force and multi-fragment dynamics for complex reactions and exotic nuclei
 - Three-cluster resonances and Borromean halo nuclei: ${}^5\text{H}$, ${}^6\text{He}$, ${}^{11}\text{Li}$, ${}^{14}\text{Be}$, ...
 - ${}^3\text{He}({}^3\text{He},2p){}^4\text{He}$ (standard solar model), ${}^3\text{H}({}^3\text{H},2n){}^4\text{He}$ (fusion research), ...

Requirements: faculties/staff, postdocs, graduate students, high-performance computing